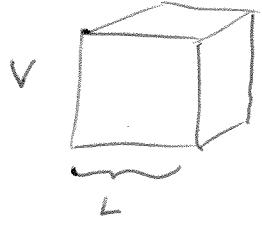


Tutorial 23/10/2014

Statistical Physics

Topic of today: IDEAL QUANTUM GAS



$$\mathcal{H} = \frac{p^2}{2m} \text{ in volume } V$$

- free, non-interacting particles
- INDISTINGUISHABLE

QM: Solution to Schrödinger equation

$$\psi_{\vec{k}}(\vec{r}) = \langle \vec{r} | \psi_{\vec{k}} \rangle = \frac{1}{\sqrt{V}} e^{i\vec{k} \cdot \vec{r}} \quad \text{with } \epsilon_{\vec{k}} = \frac{|\vec{k}|^2 \hbar^2}{2m} = \frac{p^2}{2m}$$

→ Single-particle wavefunction. For multi-particle solution, we have two options (INDIST. PTCLS)

- symmetric under ptcl exchange: $\psi(r_1, r_2) = \psi(r_2, r_1)$
- antisymmetric " " " " : $\psi(r_1, r_2) = -\psi(r_2, r_1)$

$$\Rightarrow \psi_{\{\vec{k}_i\}}(r_1, r_2, \dots, r_N) = \frac{1}{\sqrt{N!}} \sum_P \delta_P P[\psi_{\vec{k}_1}(r_1) \dots \psi_{\vec{k}_N}(r_N)]$$

↑
 Permutations over $\{\vec{k}_1, \dots, \vec{k}_N\}$

$\delta_P = \begin{cases} +1 & \text{if } P \text{ is even} \\ -1 & \text{if } P \text{ is odd} \end{cases}$
 (FERMIONS) (BOSONS)

→ $\psi_{\{\vec{k}_i\}}$ vanishes for fermions if $\vec{k}_i = \vec{k}_j$ for some $i \neq j$

⇒ For fermions & bosons,

$$\{n_{\vec{k}}\} \text{ uniquely specifies } \psi(r_1, r_2, \dots, r_N)!$$

So, in the following, it will be enough just to work with those to calculate partition function & statistical properties.

"2" options:

- a) grand-canonical ensemble (lectures):
convenient as sum in \mathbb{Z} simplifies
- b) microcanonical ensemble (will give here in tutorial)

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(Statistical Physics)

By the way: if we want to calculate a classical version, we would have a pre-factor of $\prod \frac{1}{n_k!}$ (the Gibbs 'correct counting' is implicit here)

a) Grand-canonical ensemble (see lectures)

$$Z = \sum_{\{n_k\}} \frac{\Delta}{e^{-\beta(E-\mu N)}}_{\{n_k\}} \quad \text{with} \quad E = \sum_k \epsilon_k n_k$$

$$N = \sum_k n_k$$

- FERMIONS** $n_k = 0, 1$
- BOSONS** $n_k = 0, 1, 2, \dots$

$$\Rightarrow Z = \sum_{n_1, n_2, \dots} \left[\left\{ z e^{-\beta \epsilon_1} \right\}^{n_1} \left\{ z e^{-\beta \epsilon_2} \right\}^{n_2} \dots \right] \quad (z = e^{\beta \mu})$$

this is where grand-canonical ensemble comes in handy

$$\Rightarrow \prod_k \sum_{n_k} \left(z e^{-\beta \epsilon_k} \right)^{n_k} = \begin{cases} \prod_k (1 + z e^{-\beta \epsilon_k}) & \text{FERMIONS} \\ \prod_k \frac{1}{1 - z e^{-\beta \epsilon_k}} & \text{BOSONS} \end{cases}$$

Thermodynamics

$$\Rightarrow \langle n_k \rangle = \begin{cases} \frac{1}{z^{-1} e^{\beta \epsilon_k} + 1} & \text{fermions} \\ \frac{1}{z^{-1} e^{\beta \epsilon_k} - 1} & \text{bosons} \end{cases}$$

&

$$-\beta \Omega(T, V, \mu) = \pm \sum_k \ln(1 \pm z e^{-\beta \epsilon_k}) \quad \begin{matrix} (+ \text{ fermions} \\ - \text{ bosons} \end{matrix}$$

in thermodynamic limit, k-space is packed!

$$= \pm \int g(k) dk \ln(1 \pm z e^{-\beta \epsilon_k}) \quad \text{with } g(k) dk = \frac{4\pi k^2}{(2\pi/L)^3} dk$$

$$= \pm \int g(\epsilon) d\epsilon \ln(1 \pm z e^{-\beta \epsilon}) \quad \text{with } g(\epsilon) d\epsilon = g(k) dk \text{ \& } d\epsilon = \frac{\hbar^2}{m}$$

of \vec{k} between k and $k+dk$ in 3D

Integration by parts!

$$= \pm \int d\epsilon \frac{V}{4\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2} \epsilon^{1/2} \ln(1 \pm z e^{-\beta \epsilon})$$

$$= \pm \left[\frac{2}{3} \epsilon^{3/2} \ln(1 \pm z e^{-\beta \epsilon}) \cdot \text{const.} \right]_0^\infty + \frac{2\beta}{3} \text{const.} \int_0^\infty \frac{\epsilon^{3/2}}{z^{-1} e^{\beta \epsilon} \pm 1} d\epsilon$$

$$= \frac{2\beta}{3} \int_0^\infty g(\epsilon) \epsilon / (z^{-1} e^{\beta \epsilon} \pm 1)$$

$$\Rightarrow g(\epsilon) = \frac{4\pi k^2}{(2\pi/L)^3} \cdot \frac{m}{\hbar^2 k}$$

$$= \frac{4\pi L^3 m}{(2\pi)^3 \hbar^2} \sqrt{\frac{2m\epsilon}{\hbar^2}}$$

$$= \frac{V}{4\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2} \epsilon^{1/2}$$

Similarly,

$$U = \sum_k \langle n_k \rangle \epsilon_k \approx \int d\epsilon g(\epsilon) \epsilon / (z^{-1} e^{\beta \epsilon} \pm 1) = \frac{3}{2} \Omega = \frac{3}{2} pV$$

→ looks like equipartition!

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Let us check this! What would happen in 2D, or 1D?

2D What changes is $g(E)$!

Now $g(k)dk = 2\pi k dk / (\frac{2\pi}{L})^2$



and so $g(\epsilon) = \frac{2\pi m}{(\frac{2\pi}{L})^2 \hbar^2}$ (constant!)

$$\begin{aligned} \Rightarrow \Omega &= \mp k_B T \int_0^\infty d\epsilon g(\epsilon) \ln(1 \pm z e^{-\beta\epsilon}) \\ &= \mp k_B T \left[\underbrace{\frac{2\pi m}{\hbar^2 (\frac{2\pi}{L})^2} \epsilon \ln(1 \pm z e^{-\beta\epsilon})}_0 \right]_0^\infty \pm \int_0^\infty \frac{\beta g(\epsilon) \epsilon}{z^{-1} e^{\beta\epsilon} \pm 1} d\epsilon \\ &= - \int_0^\infty g(\epsilon) \frac{\epsilon}{z^{-1} e^{\beta\epsilon} \pm 1} d\epsilon \end{aligned}$$

$$\Rightarrow U = \int_0^\infty g(\epsilon) \frac{\epsilon}{z^{-1} e^{\beta\epsilon} \pm 1} d\epsilon = -\Omega = pV \quad \checkmark$$

1D All we really have to check is whether

$$\left[G(\epsilon) \ln(1 \pm z e^{-\beta\epsilon}) \right]_0^\infty \text{ vanishes (see above).}$$

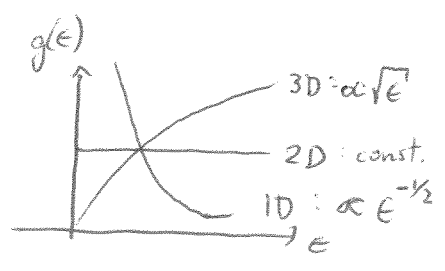
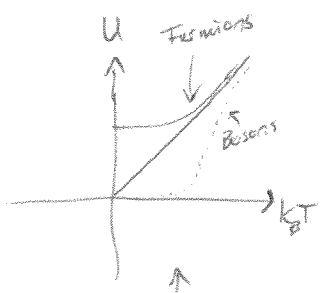
($g(\epsilon)$ integrated)

Now $g(k)dk = \frac{dk}{(2\pi/L)}$ in 1D and so $g(\epsilon) = \epsilon^{-1/2} \cdot \text{const.}$

$\Rightarrow G(\epsilon) \propto 2\epsilon^{1/2}$

$\Rightarrow [-]_0^\infty$ vanishes as required

$$\Rightarrow \Omega = - \int_0^\infty 2\epsilon \cdot g(\epsilon) / (z^{-1} e^{\beta\epsilon} \pm 1) = -2U \Rightarrow U = \frac{1}{2} pV \quad \checkmark$$



Behaviour of $g(\epsilon)$ in 1D, 2D, 3D

We shall see later that this is what happens: For large T , $pV \approx k_B T$ and classical equipartition is recovered!

(following Kerson Huang: "Introduction to statistical physics")

b) Microcanonical ensemble:

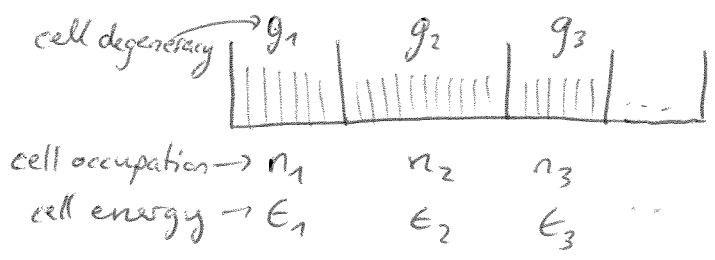
Fixed total Energy

$$E = \sum_E \epsilon_E n_E$$

" " Particle Number

$$N = \sum_E n_E$$

Simplified Picture:



(to capture the situation of k-space and associated energies)

Now, for a given set of occupations $\{n_i\}$, the number of configurations is given by

$$\Omega\{n_i\} = \prod_j \omega_j(n_j),$$

where $\omega_j(n_j)$ is the number of ways in which n_j particles can be put into g_j compartments.

FERMIONS

Each compartment can be either occupied or not.

Hence

$$\omega_j(n_j) = \frac{g_j}{n_j!(g_j - n_j)!}$$

Assume $g_j, n_j \gg 1$.

Then

$$\ln \Omega\{n_i\} = \sum_j [\ln g_j! - \ln n_j! - \ln (g_j - n_j)!]$$

$$\approx \sum_j [g_j \ln g_j - n_j \ln n_j - (g_j - n_j) \ln (g_j - n_j)]$$

To get the microcanonical ensemble, we can now maximize this quantity over variations on $\{n_i\}$ subject to the constraints of fixed N & E !

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 (Statistical Physics)

→ Use Lagrange multipliers!

$$\delta \left[\ln \Omega \{n_i\} + \alpha \sum_j n_j - \beta \sum_j \epsilon_j n_j \right] = 0$$

for independent variation of n_j .

$$\Rightarrow \sum_j \delta n_j \left[-\ln n_j + \ln (g_j - n_j) + \alpha - \beta \epsilon_j \right] = 0$$

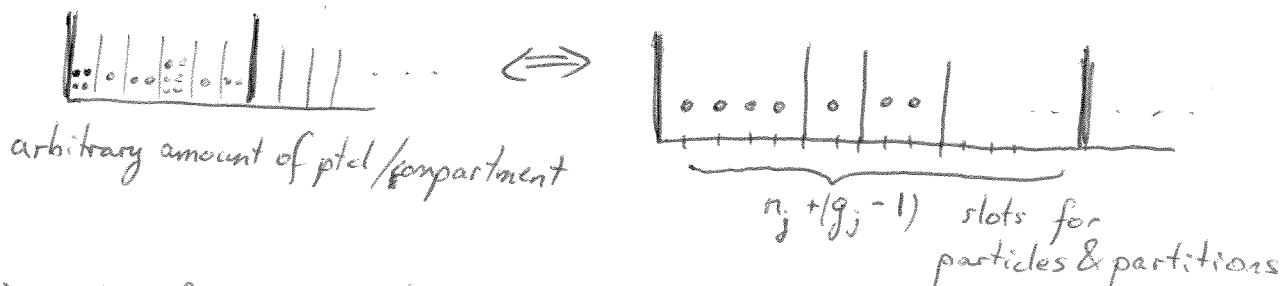
Now δn_j are independent & arbitrary, so each term independently must vanish!

$$\Rightarrow -\ln n_j + \ln (g_j - n_j) + \alpha - \beta \epsilon_j = 0$$

$$\Rightarrow n_j (e^{-\alpha + \beta \epsilon_j} + 1) = g_j$$

$$\Rightarrow \boxed{n_j = \frac{g_j}{e^{-\alpha + \beta \epsilon_j} + 1}} \quad \text{Fermi distribution}$$

BOSONS



⇒ number of ways of distributing n_j pticles:

$$\omega_j(n_j) = \frac{(n_j + g_j - 1)!}{n_j! (g_j - 1)!}$$

$$\Rightarrow \ln \Omega \{n_i\} = \prod_j \frac{(n_j + g_j - 1)!}{n_j! (g_j - 1)!} \approx \sum_j \left[(n_j + g_j) \ln (n_j + g_j) - n_j \ln n_j - g_j \ln g_j \right]$$

$g_j - 1 \approx g_j$; Stirling

Most probable distribution:

$$\sum_j \delta n_j \left[\ln (n_j + g_j - 1) - \ln n_j + \alpha - \beta \epsilon_j \right] = 0 \quad (\text{Lagrange multipliers, similar to above})$$

$$\Rightarrow \frac{n_j + g_j - 1}{n_j} = e^{\beta \epsilon_j - \alpha}$$

$$\Rightarrow \boxed{n_j = \frac{g_j - 1}{e^{\beta \epsilon_j - \alpha} - 1} \approx \frac{g_j}{e^{\beta \epsilon_j - \alpha} - 1}} \quad \text{Bose distribution}$$

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Comparison with grand-can. solution

In the microcan. ensemble, we found

$$n_j = \frac{g_j}{e^{-\alpha + \beta \epsilon_j} \pm 1} \quad \begin{matrix} + \text{ fermion} \\ - \text{ bosons} \end{matrix}$$

Because the g_j was really a tool to capture degeneracy within ϵ and $\epsilon + \Delta\epsilon$, we can now realize

$$\sum_j g_j \rightarrow \sum_{\epsilon}$$

$$\Rightarrow N = \sum_j \frac{g_j}{e^{-\alpha + \beta \epsilon_j} \pm 1} \rightarrow \sum_{\epsilon} \frac{1}{e^{-\alpha + \beta \epsilon} \pm 1}$$

In the thermodynamic limit, we thus get

$$n = \frac{N}{V} = \int \frac{d^3k}{(2\pi)^3} \frac{1}{e^{-\alpha + \beta \epsilon_k} \pm 1} = \frac{g}{\pi^2} \int_0^{\infty} dk \frac{k^2}{e^{-\alpha + \beta \epsilon_k} \pm 1}$$

Now for $T \rightarrow \infty$, this should be equal to Maxwell-Boltzmann distribution, and so we can find (for integral to be finite for $\beta \rightarrow 0$, we need $e^{-\alpha} \rightarrow \infty$. Hence $n_{\epsilon} \xrightarrow{\beta \rightarrow 0} e^{\alpha - \beta \epsilon_{\epsilon}}$)

$$\beta = \frac{1}{k_B T}$$

Also, we can define

$$\mu = \frac{\alpha}{\beta}; \quad z = e^{\beta \mu}$$

Now we see that

$$n_{\epsilon} = \frac{1}{z^{-1} e^{\beta \epsilon} \pm 1}$$

which corresponds to $\langle n_{\epsilon} \rangle$ from the grand-canonical ensemble!

Final Remark: classical particles! (see box on page 2)

For $e^{-\beta(\epsilon_k - \mu)} \ll 1$ (high temperature / low density)

$$\Omega = -\beta^{-1} \ln Z \approx -\beta^{-1} \sum_{\mathbf{k}} e^{-\beta(\epsilon_{\mathbf{k}} - \mu)} = -\frac{z}{\beta} \frac{V}{(2\pi)^3} \int d^3k e^{-\beta k^2 / 2m} = -\frac{zV}{\beta \lambda^3}$$

with $\lambda = \frac{h}{\sqrt{2\pi m k T}}$ (thermal wavelength).

Now classically, we calculate (canonical) $Z = \sum_{N=0}^{\infty} \frac{1}{N!} e^{\beta \mu N} \left(\frac{V}{\lambda^3}\right)^N = e^{zV/\lambda^3}$

or equivalently:

$$Z = \sum_{n_1, n_2, \dots} \left[\frac{1}{n_1!} \left\{ z e^{-\beta \epsilon_1} \right\}^{n_1} \frac{1}{n_2!} \left\{ z e^{-\beta \epsilon_2} \right\}^{n_2} \dots \right] = \sum_{N=0}^{\infty} z^N \frac{1}{N!} \left\{ \sum_{\epsilon} e^{-\beta \epsilon} \right\}^N$$

(SEE BOX ON PAGE 2!)
"correct counting" \uparrow