ETH Zurich, HS12

2 Classical Free Scalar Field

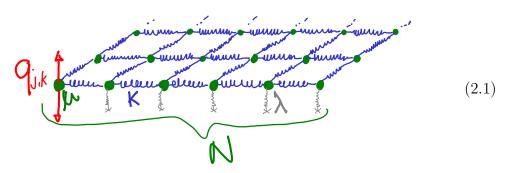
In the following we shall discuss one of the simplest field theory models, the classical non-interacting relativistic scalar field.

2.1 Spring Lattice

Before considering field, start with an approximation we can certainly handle: lattice.

Consider an atomic lattice:

- 1D or 2D cubic lattice,
- atoms are coupled to neighbours by springs, 1
- atoms are coupled to rest position by springs,
- atoms can move only orthogonally to lattice (transverse),
- boundaries: periodic identification.



model parameters and variables:

- lattice separation r,
- number of atoms N (in each direction),
- mass of each atom μ ,
- lattice spring constant κ ,
- return spring constant λ ,
- shift orthogonal to lattice $q_{i,k}$

Lagrangian Formulation. Lagrange function

$$L = L_{\rm kin} - V_{\rm lat} - V_{\rm rest} \tag{2.2}$$

¹Springs are useful approximations because they model first deviation from rest position; always applies to small excitations.

Standard non-relativistic kinetic terms

$$L_{\rm kin} = \frac{1}{2}\mu \sum_{i,j=1}^{N} \dot{q}_{i,j}^{2}.$$
 (2.3)

Potential for springs between atoms (ignore x, y-potential)²

$$V_{\text{lat}} = \frac{1}{2}\kappa \sum_{i,j=1}^{N} (q_{i-1,j} - q_{i,j})^2 + \frac{1}{2}\kappa \sum_{i,j=1}^{N} (q_{i,j-1} - q_{i,j})^2.$$
 (2.4)

Some spring potential to drive atoms back to rest position

$$V_{\text{rest}} = \frac{1}{2}\lambda \sum_{i,j=1}^{N} q_{i,j}^{2}.$$
 (2.5)

Quadratic in q's: bunch of coupled HO's. Equations of motion

$$\mu \ddot{q}_{i,j} = -\kappa (q_{i-1,j} - 2q_{i,j} + q_{i+1,j}) - \kappa (q_{i,j-1} - 2q_{i,j} + q_{i,j+1}) + \lambda q_{i,j} = 0.$$
(2.6)

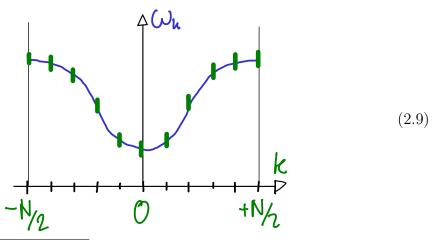
Note: spatially homogeneous equations. Use discrete Fourier transform to solve (respect periodicity)

$$q_{i,j}(t) = \frac{1}{N^2} \sum_{k,l=1}^{N} \frac{c_{k,l}}{\sqrt{2\mu\omega_{k,l}}} \exp\left(-\frac{2\pi i}{N}(ki+lj) - i\omega_{k,l}t\right) + \frac{1}{N^2} \sum_{k,l=1}^{N} \frac{c_{k,l}^*}{\sqrt{2\mu\omega_{k,l}}} \exp\left(\frac{2\pi i}{N}(ki+lj) + i\omega_{k,l}t\right)$$
(2.7)

Used freedom to define coefficients $c_{k,l}$ to introduce prefactors. Complex conjugate coefficients $c_{k,l}^*$ in second term ensure reality of $q_{i,j}$. Note: $c_{k,l}^*$ represents mode of opposite momentum and energy w.r.t. $c_{k,l}$!

E.o.m. translates to dispersion relation:

$$\mu\omega_{k,l}^2 = \lambda + 4\kappa\sin^2\frac{\pi k}{N} + 4\kappa\sin^2\frac{\pi l}{N}.$$
 (2.8)



Can ignore due to $d^2 = d_x^2 + d_y^2 + d_z^2$, shifts potential by an irrelevant constant. Moreover longitudinal and transverse excitations decouple.

Hamiltonian Formulation. Define momenta

$$p_{i,j} := \frac{\partial L}{\partial \dot{q}_{i,j}} = \mu \dot{q}_{i,j}. \tag{2.10}$$

Derive Hamiltonian function as Legendre transform of L

$$H = H_{\text{kin}} + V_{\text{lat}} + V_{\text{rest}}$$
 with $H_{\text{kin}} = \frac{1}{2\mu} \sum_{i,j=1}^{N} p_{i,j}^2$. (2.11)

Then define canonical Poisson brackets

$$\{f,g\} := \sum_{i,j=1}^{N} \left(\frac{\partial f}{\partial q_{i,j}} \frac{\partial g}{\partial p_{i,j}} - \frac{\partial f}{\partial p_{i,j}} \frac{\partial g}{\partial q_{i,j}} \right).$$
 (2.12)

In other words $\{q_{i,j}, p_{j,k}\} = \delta_{i,k}\delta_{j,l}$ and $\{q, q\} = \{p, p\} = 0$.

Fourier Modes. Introduce new complex variables (Fourier transform)

$$a_{k,l} = \frac{1}{\sqrt{2\mu\omega_{k,l}}} \sum_{i,j=1}^{N} \exp\left(\frac{2\pi i}{N}(ki+lj)\right) \left(\mu\omega_{k,l}q_{i,j} + ip_{i,j}\right).$$
 (2.13)

Transformed Hamiltonian is very simple

$$H = \frac{1}{N^2} \sum_{k,l=1}^{N} \omega_{k,l} a_{k,l}^* a_{k,l}. \tag{2.14}$$

Poisson brackets for new variables are simple, too

$$\{a_{i,j}, a_{k,l}^*\} = -i N^2 \delta_{i,k} \delta_{j,l},$$

$$\{a_{i,j}, a_{k,l}\} = \{a_{i,j}^*, a_{k,l}^*\} = 0.$$
 (2.15)

Can convince oneself that equations of motion hold

$$\dot{a}_{k,l} = -\{H, a_{k,l}\} = -i\omega_{k,l}a_{k,l},
\dot{a}_{k,l}^* = -\{H, a_{k,l}^*\} = +i\omega_{k,l}a_{k,l}^*.$$
(2.16)

and solved by above Lagrangian solution

$$a_{k,l}(t) = c_{k,l} \exp(-i\omega_{k,l}t), \qquad a_{k,l}^*(t) = c_{k,l}^* \exp(+i\omega_{k,l}t).$$
 (2.17)

2.2 Continuum Limit

Now turn this spring lattice into a smooth field φ :

- send number of sites $N \to \infty$.
- box of size L in all directions; lattice separation $r = L/N \to 0$.
- positions x = ir = iL/N,
- field $q_{i,...} = \varphi(\vec{x})$,
- generalise to d spatial dimensions, e.g. d = 1, 2, 3.

Some useful rules

$$\sum_{i=1}^{N} \to \frac{1}{r} \int dx, \qquad q_i - q_{i-1} \to r(\partial \varphi)$$
 (2.18)

Lagrangian Formulation. Substitute this in Lagrangian

$$L \to \int d^d \vec{x} \left(\frac{\mu}{2r^d} \dot{\varphi}^2 - \frac{\kappa}{2r^{d-2}} (\vec{\partial}\varphi)^2 - \frac{\lambda}{2r^d} \varphi^2 \right). \tag{2.19}$$

Diverges as $r \to 0$, but can rescale parameters. Suitable rescalings

$$\mu = r^d \bar{\mu}, \quad \kappa = r^{d-2} \bar{\kappa}, \quad \lambda = r^d \bar{\lambda},$$
 (2.20)

Parameters become densities. Lagrangian functional

$$L[\varphi, \dot{\varphi}](t) = \int d^d \vec{x} \left(\frac{1}{2} \bar{\mu} \dot{\varphi}^2 - \frac{1}{2} \bar{\kappa} (\vec{\partial} \varphi)^2 - \frac{1}{2} \bar{\lambda} \varphi^2 \right). \tag{2.21}$$

Can furthermore rescale field $\varphi = \bar{\kappa}^{-1/2} \phi$

$$L[\phi, \dot{\phi}](t) = \int d^d \vec{x} \left(\frac{1}{2} \bar{\mu} \bar{\kappa}^{-1} \dot{\phi}^2 - \frac{1}{2} (\vec{\partial} \phi)^2 - \frac{1}{2} \bar{\lambda} \bar{\kappa}^{-1} \phi^2 \right). \tag{2.22}$$

Derive e.o.m.: start with action functional $S[\phi]$

$$S[\phi] = \int dt L[\phi](t) = \int dt d^d \vec{x} \mathcal{L}(\phi(\vec{x}, t), \partial_i \phi(\vec{x}, t), \dot{\phi}(\vec{x}, t)); \qquad (2.23)$$

useful to express (homogeneous) Lagrange functional $L[\phi](t)$ through Lagrangian density $\mathcal{L}(\phi, \vec{\partial}\phi, \dot{\phi})$ (the Lagrangian).

Vary action functional (discard boundary terms)

$$\delta S[\phi] = \int dt \, d^d \vec{x} \, \delta \phi \left(\frac{\partial \mathcal{L}}{\partial \phi} - \frac{\partial}{\partial x^i} \, \frac{\partial \mathcal{L}}{\partial (\partial_i \phi)} - \frac{d}{dt} \, \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) + \dots \stackrel{!}{=} 0.$$
 (2.24)

Write general Euler–Lagrange equation for fields

$$\frac{\partial \mathcal{L}}{\partial \phi}(\vec{x}, t) - \frac{\partial}{\partial x^i} \frac{\partial \mathcal{L}}{\partial (\partial_i \phi)}(\vec{x}, t) - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\phi}}(\vec{x}, t) = 0.$$
 (2.25)

In our case

$$-\bar{\mu}\bar{\kappa}^{-1}\ddot{\phi} + \vec{\partial}^{\,2}\phi - \bar{\lambda}\bar{\kappa}^{-1}\phi = 0; \tag{2.26}$$

agrees with continuum limit of discrete e.o.m.. Now denote as speed of light c and mass m

$$\bar{\mu}\bar{\kappa}^{-1} = c^{-2} = 1, \qquad \bar{\lambda}\bar{\kappa}^{-1} = m^2,$$
 (2.27)

to discover Klein-Gordon equation (set c=1)

$$-c^{-2}\ddot{\phi} + \vec{\partial}^{\,2}\phi - m^2\phi = 0. \tag{2.28}$$

Plane Wave Solutions. Consider solutions on infinite space and time. Homogeneous equation solved by Fourier transformation

$$\phi(\vec{x},t) = \int \frac{d^d \vec{p}}{(2\pi)^d 2e(\vec{p})} \alpha(\vec{p}) \exp(-i\vec{p}\cdot\vec{x} - ie(\vec{p})t)$$

$$+ \int \frac{d^d \vec{p}}{(2\pi)^d 2e(\vec{p})} \alpha^*(\vec{p}) \exp(+i\vec{p}\cdot\vec{x} + ie(\vec{p})t)$$
(2.29)

with (positive) energy $e(\vec{p})$ on mass shell (often called $\omega(\vec{p})$)

$$e(\vec{p}) = \sqrt{\vec{p}^2 + m^2}. (2.30)$$

Agrees with discrete solution identifying momenta as

$$p = 2\pi k/L, \qquad \sum_{k=1}^{N} \to \frac{L}{2\pi} \int dp, \qquad c_{k,\dots} = \frac{\alpha(\vec{p})}{\sqrt{2e(\vec{p})r^d}}.$$
 (2.31)

Some remarks on factors and conventions:

- Fourier transforms on \mathbb{R} produce factors of 2π , need to be put somewhere. Convention to associate $(2\pi)^{-1}$ to every dp: $dp := dp/2\pi$. No factors of 2π for dx. No factors of 2π in exponent.
- Combination $d^d\vec{p}/2e(\vec{p})$ is a useful combination: Relativistic covariance. Reason for conversion factor in $c_{k,...}$.

2.3 Relativistic Covariance

The Klein–Gordon equation can be written manifestly relativistically 3 4

$$-\partial^{\mu}\partial_{\mu}\phi + m^{2}\phi = -\partial^{2}\phi + m^{2}\phi = 0.$$
 (2.32)

Also Lagrangian and action manifestly relativistic

$$\mathcal{L} = -\frac{1}{2}(\partial\phi)^2 - \frac{1}{2}m^2\phi^2, \qquad S = \int d^D x \,\mathcal{L}. \tag{2.33}$$

To understand the relativistic behaviour of the solution, consider integration over a mass shell $p^2 + m^2 = 0$

$$\int d^{D}p \, \delta(p^{2} + m^{2}) \theta(p^{0}) \, f(p)$$

$$= \int d^{d}\vec{p} \, de \, \delta(-e^{2} + \vec{p}^{2} + m^{2}) \theta(e) \, f(e, \vec{p})$$

$$= \int \frac{d^{d}\vec{p} \, de}{2e} \, \delta(e - \sqrt{\vec{p}^{2} + m^{2}}) \theta(e) \, f(e, \vec{p})$$

$$= \int \frac{d^{d}\vec{p}}{2e(\vec{p})} \, f(e(\vec{p}), \vec{p})$$
(2.34)

 $^{^{3}\}partial^{2}$ often written as D'Alembertian \square .

 $^{^4}$ Signature of spacetime is -+++!

Solution is just Fourier transform

$$\phi(x) = \int \frac{d^D p}{(2\pi)^D} \,\phi(p) \exp(-ip_\mu x^\mu) \tag{2.35}$$

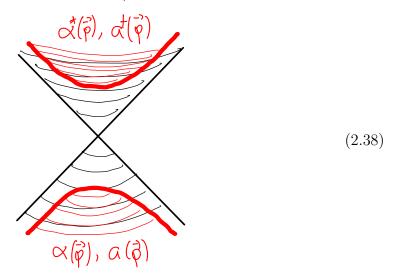
with momentum space field defined on shell only

$$\phi(p) = 2\pi\delta(p^2 + m^2) (\theta(p_0)\alpha(\vec{p}) + \theta(-p_0)\alpha^*(-\vec{p})).$$
 (2.36)

Momentum space e.o.m. obviously satisfied

$$(p^2 + m^2)\phi(p) = 0. (2.37)$$

 $\alpha(\vec{p})$ and $\alpha^*(\vec{p})$ define amplitudes on forward/backward mass shells.



Note that $\phi(p)$ obeys reality condition (from $\phi(x)^* = \phi(x)$)

$$\phi(p)^* = \phi(-p). \tag{2.39}$$

2.4 Hamiltonian Field Theory

Now that we have a nice relativistic formulation for the Klein–Gordon field $\phi(x)$, let's separate space from time. ⁵

Position Space. Define momentum π (field) conjugate to ϕ

$$\pi(\vec{x},t) = \frac{\delta L}{\delta \dot{\phi}(\vec{x})}(t) = \frac{\partial \mathcal{L}}{\partial \dot{\phi}}(\vec{x},t) = \dot{\phi}(\vec{x},t). \tag{2.40}$$

Determine Hamiltonian

$$H[\phi, \pi] = \int d^{d}\vec{x} \,\pi \dot{\varphi} - L[\phi, \dot{\phi}]$$

$$= \int d^{d}\vec{x} \,\left(\frac{1}{2}\pi^{2} + \frac{1}{2}(\vec{\partial}\phi)^{2} + \frac{1}{2}m^{2}\phi^{2}\right). \tag{2.41}$$

 $^{^5{\}mbox{Formalism}}$ breaks relativistic invariance, physics remains relativistic.

Not relativistically covariant, not designed to be.⁶

Define Poisson brackets for phase space functionals f, g

$$\{f,g\} = \int d^d \vec{x} \left(\frac{\delta f}{\delta \phi(\vec{x})} \frac{\delta g}{\delta \pi(\vec{x})} - \frac{\delta f}{\delta \pi(\vec{x})} \frac{\delta g}{\delta \phi(\vec{x})} \right). \tag{2.42}$$

Poisson brackets of fields yield delta-functions⁷

$$\{\phi(\vec{x}), \pi(\vec{y})\} = \int d^d \vec{z} \, \delta^d(\vec{x} - \vec{z}) \delta^d(\vec{y} - \vec{z}) = \delta^d(\vec{x} - \vec{z}). \tag{2.43}$$

Momentum Space. Now introduce momentum modes⁸

$$a(\vec{p}) = \int d^d \vec{x} \exp(i\vec{p}\cdot\vec{x}) \left(e(\vec{p})\phi(\vec{x}) + i\pi(\vec{x})\right), \tag{2.44}$$

and inverse Fourier transformation

$$\phi(\vec{x}) = \int \frac{d^{d}\vec{p}}{(2\pi)^{d} 2e(\vec{p})} a(\vec{p}) \exp(-i\vec{p}\cdot\vec{x}) + \int \frac{d^{d}\vec{p}}{(2\pi)^{d} 2e(\vec{p})} a^{*}(\vec{p}) \exp(+i\vec{p}\cdot\vec{x}), \pi(\vec{x}) = -\frac{i}{2} \int \frac{d^{d}\vec{p}}{(2\pi)^{d}} a(\vec{p}) \exp(-i\vec{p}\cdot\vec{x}) + \frac{i}{2} \int \frac{d^{d}\vec{p}}{(2\pi)^{d}} a^{*}(\vec{p}) \exp(+i\vec{p}\cdot\vec{x}).$$
(2.45)

Compute Poisson brackets for Fourier modes⁹ 10

$$\{a(\vec{p}), a^*(\vec{q})\} = -i \, 2e(\vec{p}) \, (2\pi)^d \delta^d(\vec{p} - \vec{q}). \tag{2.46}$$

In other words

$$\{f,g\} = -i(2\pi)^d \int d^d\vec{p} \ 2e(\vec{p}) \left(\frac{\delta f}{\delta a(\vec{p})} \frac{\delta g}{\delta a^*(\vec{p})} - \frac{\delta f}{\delta a^*(\vec{p})} \frac{\delta g}{\delta a(\vec{p})} \right). \tag{2.47}$$

Hamiltonian translates to

$$H = \frac{1}{2} \int \frac{d^d \vec{p}}{(2\pi)^d} a^*(\vec{p}) a(\vec{p}) = \int \frac{d^d \vec{p}}{(2\pi)^d 2e(\vec{p})} e(\vec{p}) a^*(\vec{p}) a(\vec{p}). \tag{2.48}$$

Hence e.o.m. for field oscillator

$$\dot{a}(\vec{p}) = -\{H, a(\vec{p})\} = -ie(\vec{p})a(\vec{q}),
\dot{a}^*(\vec{p}) = -\{H, a^*(\vec{p})\} = +ie(\vec{p})a^*(\vec{q}).$$
(2.49)

One HO for every momentum. Solution

$$a(\vec{p},t) = \alpha(\vec{p}) \exp(-ie(\vec{p})t), \quad a^*(\vec{p},t) = \alpha^*(\vec{p}) \exp(+ie(\vec{p})t). \tag{2.50}$$

⁶Hamiltonian governs *time* translation.

⁷Formula for variation $\delta\phi(\vec{x})/\delta\phi(\vec{z}) = \delta^d(\vec{x} - \vec{z})$.

⁸Have some additional factors compared to some literature.

 $^{^{9} \}text{Conventional } 2\pi$ for delta-function in momentum space.

¹⁰Factor $2e(\vec{p})$ appropriate relativistic measure for mass shell.