

**Exercise 4.1 Local equilibrium state of a gas in a periodic potential**

We consider a gas of  $N$  particles trapped in a box,  $\vec{r} \in V = [0, L]^3$ , in the presence of a conservative force  $\vec{F}(\vec{r}) = -\nabla\mathcal{V}(\vec{r})$  originating from a periodic potential in the  $x$ -direction,

$$\mathcal{V}(\vec{r}) = \mathcal{V}_0 \cos(2\pi kx/L), \quad k \in \mathbb{N}. \quad (1)$$

For the distribution function in equilibrium, we make the following ansatz; we generalize the free-particle (Maxwell-Boltzmann) distribution function by taking into account a local density  $n(\vec{r})$ , i.e.,

$$f_0(\vec{r}, \vec{p}) = \frac{n(\vec{r})}{(2\pi m k_B T)^{3/2}} e^{-\beta p^2/2m}, \quad \beta = \frac{1}{k_B T}. \quad (2)$$

- Find the local density  $n(\vec{r})$ . Discuss the limits  $\beta\mathcal{V}_0 \ll 1$  and  $\beta\mathcal{V}_0 \gg 1$ .
- Determine the internal energy  $U$  and the specific heat  $C_V$ . Discuss these expressions in the limits  $\beta\mathcal{V}_0 \ll 1$  and  $\beta\mathcal{V}_0 \gg 1$ .
- Calculate the entropy  $S = S(T, V, N)$ .

*Hints:* The integral representation and the series expansion of the modified Bessel functions of the first kind for  $n \in \mathbb{Z}$  are

$$I_n(z) = \frac{1}{\pi} \int_0^\pi d\theta e^{z \cos \theta} \cos(n\theta) = \left(\frac{z}{2}\right)^n \sum_{k \geq 0} \frac{(z^2/4)^k}{k!(n+k)!}.$$

The asymptotic behavior for  $z \rightarrow \infty$  is

$$I_n(z) \sim \frac{e^z}{\sqrt{2\pi z}} \left[ 1 - \frac{4n^2 - 1}{8z} + \dots \right].$$

Furthermore, the relation  $I'_0(z) = I_1(z)$  holds.

**Exercise 4.2 Distribution function for relativistic massive particles**

Consider a system of relativistic massive particles with vanishing drift velocity  $\vec{v}_{\text{drift}} = 0$ . The energy-momentum relation of the particles writes

$$E(\vec{p}) = \sqrt{p^2 c^2 + m^2 c^4}, \quad (3)$$

where  $m$  is the particle mass and  $c$  the speed of light.

- Find the equilibrium distribution function  $f_0(\vec{p})$  and show that in the limit  $k_B T \ll mc^2$ , the classical Maxwell-Boltzmann distribution function is recovered. Calculate the internal energy  $U$  and determine the first relativistic corrections to this expression as well as to the specific heat  $C_V$ .

- b) Calculate the pressure  $P$  for this system and convince yourself that in the limit  $k_B T/mc^2 \ll 1$ , it is consistent with the pressure of a dilute ideal gas as derived in the lecture notes.
- c) Take into account a finite drift velocity  $\vec{v}_{\text{drift}}$ . How does the distribution function change? Does the temperature depend on the drift velocity? Calculate the average momentum  $\langle \vec{p} \rangle$  for this system.

*Hints:* The integral representation of the modified Bessel functions of the second kind is

$$K_n(z) = \int_0^\infty dy e^{-z \cosh y} \cosh(ny),$$

Its asymptotic behavior for  $z \rightarrow \infty$  is

$$K_n(z) \sim \sqrt{\frac{\pi}{2z}} e^{-z} \left[ 1 + \frac{4n^2 - 1}{8z} + \frac{(4n^2 - 1)(4n^2 - 9)}{2!(8z)^2} + \frac{(4n^2 - 1)(4n^2 - 9)(4n^2 - 25)}{3!(8z)^3} + \dots \right].$$

**Office Hours:** Monday, October 15th, 8–10 AM (Michael Walter, HIT K 31.5).