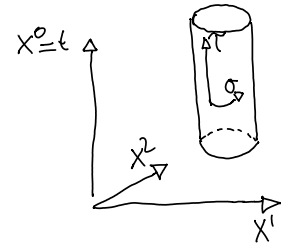


3 Classical Bosonic String

Mechanics of 1D extended object without inner structure.

- Worldsheet coordinates: $\xi^\alpha = (\tau, \sigma)$, time τ , space σ .
- Embedding coordinates $X^\mu(\xi)$.
- D -dimensional embedding Minkowski space, metric $\eta_{\mu\nu}$.



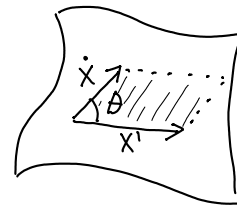
3.1 Nambu–Goto Action

Generalise insights from relativistic point particle:

- worldline \rightarrow worldsheet.
- action = proper time \simeq “length” \rightarrow “area”.

Area and Action. Wick rotation $t = iw$. Area of 2D euclidean surface:

$$\begin{aligned} dA &= d\tau d\sigma |X'| |\dot{X}| |\sin \theta| \\ &= d\tau d\sigma \sqrt{X'^2 \dot{X}^2 \sin^2 \theta} \\ &= d\tau d\sigma \sqrt{X'^2 \dot{X}^2 - (X' \cdot \dot{X})^2} \\ &= d^2\xi \sqrt{\det g}, \end{aligned}$$



induced worldsheet metric $g_{\alpha\beta} = \eta_{\mu\nu} \partial_\alpha X^\mu \partial_\beta X^\nu$ (pull back).

Wick rotate back; string action (Nambu–Goto)

$$S = -\frac{1}{2\pi\kappa^2} A = -\frac{1}{2\pi\kappa^2} \int d^2\xi \sqrt{-\det g}.$$

Symmetries of action:

- **Lorentz:** scalar products of Lorentz vectors X .
- **Poincaré:** X only through ∂X .
- **Worldsheet Diffeomorphisms:** density $d^2\xi \sqrt{-\det g}$ invariant under reparametrisations $\xi \mapsto \xi'(\xi)$.

Tension. What is parameter κ ?

Fundamental string length scale \rightarrow quantum string.



Consider potential U : Time slice of action/area. Slice of length L : $U \sim L/\kappa^2$.
Constant force $U' = 1/2\pi\kappa^2 = T$ is string tension. String not a spring or rubber band!

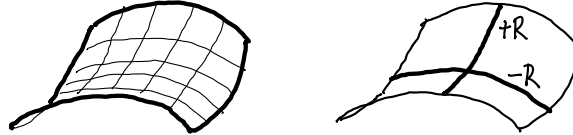
Equations of Motion. Vary action.

Use variation of determinant $\delta \det g = \det g g^{\alpha\beta} \delta g_{\alpha\beta}$.

$$\partial_\alpha (\sqrt{-\det g} g^{\alpha\beta} \partial_\beta X^\mu) = 0.$$

Highly non-linear equations (g contains X)! How to solve? How to deal with?

Geometrically: stationary action; minimal area surface. Static soap films! Mean curvature zero. Saddle point everywhere, equal/opposite sectional curvatures.



3.2 Polyakov Action

Complication from non-linear e.o.m.. As for point particle, there is a polynomial action with additional dynamical worldsheet metric $g_{\alpha\beta}$:

$$S = -\frac{1}{2\pi\kappa^2} \int d^2\xi \sqrt{-\det g} \frac{1}{2} g^{\alpha\beta} (\partial_\alpha X) \cdot (\partial_\beta X).$$

E.o.m. for X as above; e.o.m. for worldsheet metric g :

$$(\partial_\alpha X) \cdot (\partial_\beta X) = \frac{1}{2} g_{\alpha\beta} g^{\gamma\delta} (\partial_\gamma X) \cdot (\partial_\delta X).$$

Solution fixes induced metric (up to local scale f)

$$g_{\alpha\beta} = f(\xi) (\partial_\alpha X) \cdot (\partial_\beta X).$$

Arbitrary scale $f(\xi)$ cancels in action and all e.o.m.. New redundancy: **Weyl invariance** $g_{\alpha\beta}(\xi) \mapsto f(\xi) g_{\alpha\beta}(\xi)$.

3.3 Conformal Gauge

E.o.m. for X linear, coupling to g makes non-linear.

Use gauge freedom: demand conformally flat metric.

$$g_{\alpha\beta}(\xi) = f(\xi) \eta_{\alpha\beta}.$$

Amounts to two equations $g_{\tau\sigma}(\xi) = 0$, $g_{\tau\tau}(\xi) = -g_{\sigma\sigma}(\xi)$.

- Fixes almost all diffeomorphisms.
- Conformal transformations remain. Diffeomorphisms preserving metric up to scale.
- May further set $f = 1$ by means of fixing Weyl.

Action describes D free massless scalar particles in 2D

$$S = -\frac{1}{2\pi\kappa^2} \int d^2\xi \frac{1}{2} \eta^{\alpha\beta} (\partial_\alpha X) \cdot (\partial_\beta X).$$

E.o.m. simply $\partial^2 X^\mu = 0$ (harmonic wave equation)

$$\ddot{X} = X''.$$

Do not forget e.o.m. for worldsheet metric

$$T_{\alpha\beta} := (\partial_\alpha X) \cdot (\partial_\beta X) - \frac{1}{2} \eta_{\alpha\beta} \eta^{\gamma\delta} (\partial_\gamma X) \cdot (\partial_\delta X) = 0.$$

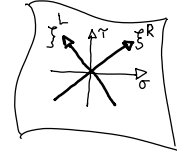
$T_{\alpha\beta}$ is energy-momentum tensor for D scalars. Trace absent $T^\alpha_\alpha = 0$ by construction (Weyl/conformal). Two remaining e.o.m. become **Virasoro constraints**

$$\dot{X} \cdot X' = 0, \quad \dot{X} \cdot \dot{X} + X' \cdot X' = 0.$$

Forbids longitudinal waves along the worldsheet, no structure!

Conservation of $T_{\alpha\beta}$: impose constraints only on time slice.

3.4 Solution on the Light Cone



Solve harmonic wave equation. Light cone coordinates $\xi^{L/R}$ useful:

$$\xi^{L/R} = \tau \mp \sigma, \quad \partial_{L/R} = \frac{1}{2}(\partial_\tau \mp \partial_\sigma),$$

new worldsheet metric

$$d^2s = -d\tau^2 + d\sigma^2 = -d\xi^L d\xi^R.$$

Now e.o.m. and Virasoro constraints read

$$\partial_L \partial_R X^\mu = 0, \quad (\partial_{L/R} X)^2 = 0.$$

First equation solved by simple separation of variables

$$X^\mu(\xi^L, \xi^R) = X_L^\mu(\xi^L) + X_R^\mu(\xi^R).$$

D left-movers X_L plus D right-movers X_R . Virasoro constraints $(\partial X_{R,L})^2 = 0$ remove 1 left/right-mover. Two reparametrisations left:

- **conformal transformations**

$$\xi^R \mapsto \xi'^R(\xi^R), \quad \xi^L \mapsto \xi'^L(\xi^L).$$

2D case special: infinitely many transformations. removes another 1 left/right-mover.

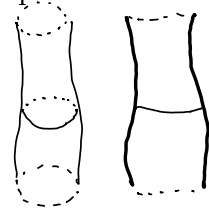
- **constant shift** between X_L^μ and X_R^μ .

$(D - 2)$ left/right-movers remain (transverse).

3.5 Closed String Modes

So far worldsheet infinitely extended in space and time; want finite spatial extent.

- **Closed String:** circular topology.
Identify $\sigma \equiv \sigma + 2\pi$ (other choices possible).
- **Open String:** interval topology.
Boundary conditions at $\sigma = 0, \pi$ (later).



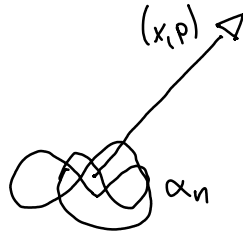
Periodic function $X \rightarrow$ Fourier decomposition

$$X_{L/R}^\mu = \frac{1}{2}x^\mu + \frac{1}{2}\kappa^2 p^\mu \xi^{L/R} + \sum_{n \neq 0} \frac{i\kappa}{\sqrt{2n}} \alpha_n^{L/R, \mu} \exp(-in\xi^{L/R}).$$

- Coefficients $i\kappa/\sqrt{2n}$ chosen for later convenience.
- Linear dependence in $\xi^{L/R}$ okay: $X^\mu = x^\mu + \kappa^2 p^\mu \tau + \dots$
- Reality of X : complex conjugate $\alpha_{-n} = (\alpha_n)^*$.

Two kinds of parameters for solutions

- Centre of mass motion x, p (conjugate variables $\rightarrow \kappa^2$).
- String modes $\alpha_n^{L/R, \mu}$ (left/right movers).



Consider Virasoro constraints, substitute modes:

$$(\partial_{L,R} X_{L/R}^\mu)^2 = \kappa^2 \sum_n L_n^{L/R} \exp(-in\xi^{L/R}) \stackrel{!}{=} 0$$

with Virasoro modes (drop L/R index)

$$L_n := \frac{1}{2} \sum_m \alpha_{n-m} \cdot \alpha_m \stackrel{!}{=} 0.$$

Note that $\alpha_0^L = \alpha_0^R = \kappa p / \sqrt{2}$. $L_0 = 0$ constraint fixes string mass

$$p^2 = -M^2, \quad M^2 = \frac{4}{\kappa^2} \sum_{m=1}^{\infty} \alpha_{-m} \cdot \alpha_m$$

All Virasoro constraints $L_n = 0$ conserved by e.o.m.

$$\dot{L}_n = inL_n.$$

Need to impose on initial data only.

String mass depends on mode amplitudes.

- No modes excited: massless point particle.
- Few/small excitations: light tiny particle.
- Many/large excitations: big massive object.



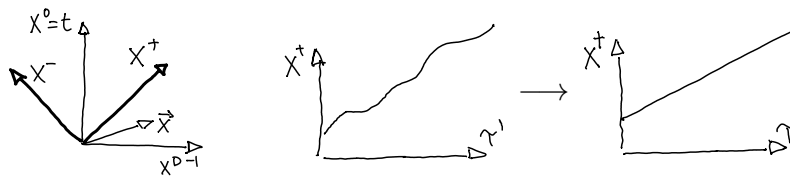
Note: time-like modes α^0 contribute negative M^2 . Tachyons excluded by Virasoro constraints.

Virasoro constraints non-linear, complicated.

3.6 Light Cone Gauge

Use conformal symmetry to solve Virasoro constraints.

Gauge Fixing. Introduce as well light cone coordinates for spacetime $X^\pm = X^0 \pm X^{D-1}$. \vec{X} denotes components $1 \dots (D-2)$. Gauge fix!



$$X_{L/R}^+ = x_{L/R}^+ + \frac{1}{2}\kappa^2 p_{L/R}^+ \xi^{L/R}.$$

Virasoro constraint $(\partial \vec{X}_{L/R})^2 - \kappa^2 p_{L/R}^+ \partial X_{L/R}^- = 0$ solved by

$$X_{L/R}^-(\xi) = \frac{1}{\kappa^2 p_{L/R}^+} \int^\xi d\xi' (\partial \vec{X}_{L/R}(\xi'))^2.$$

Solution: $2(D-2)$ arbitrary functions $\vec{X}_{L/R}(\xi^{L/R})$.

Periodicity. All functions $\vec{X}(\xi)$ periodic! Furthermore: periodicity of X^+ and X^- requires

$$p_L^+ = p_R^+ = p^+, \quad \int_0^{2\pi} d\xi ((\partial \vec{X}_R)^2 - (\partial \vec{X}_L)^2) = 0.$$

Residual gauge freedom: Constant shift $\Delta X_R^\mu(\xi) = -\Delta X_L^\mu(\xi)$. Corresponds to above residual constraints.

String Modes. Impose gauge fixing on modes ($n \neq 0$)

$$\alpha_n^+ = 0, \quad \alpha_n^- = \frac{1}{\alpha_0^+} \sum_m \vec{\alpha}_{n-m} \cdot \vec{\alpha}_m.$$

Periodicity requires $\vec{\alpha}_0^L = \vec{\alpha}_0^R$, $\alpha_0^{R,+} = \alpha_0^{L,+}$ and for α_0^-

$$\sum_{m=1}^{\infty} \vec{\alpha}_{-m}^L \cdot \vec{\alpha}_m^L = \sum_{m=1}^{\infty} \vec{\alpha}_{-m}^R \cdot \vec{\alpha}_m^R$$

Resulting mass manifestly positive:

$$M^2 = \frac{4}{\kappa^2} \sum_{m=1}^{\infty} \vec{\alpha}_{-m} \cdot \vec{\alpha}_m = \frac{4}{\kappa^2} \sum_{m=1}^{\infty} |\vec{\alpha}_m|^2$$

Benefits: positivity, almost all constraints gone. Drawback: Lorentz symmetry not manifest.